

Introduction to Cosmology



Lecture 16

1.5'

Table 1.1: COSMIC INVENTORY

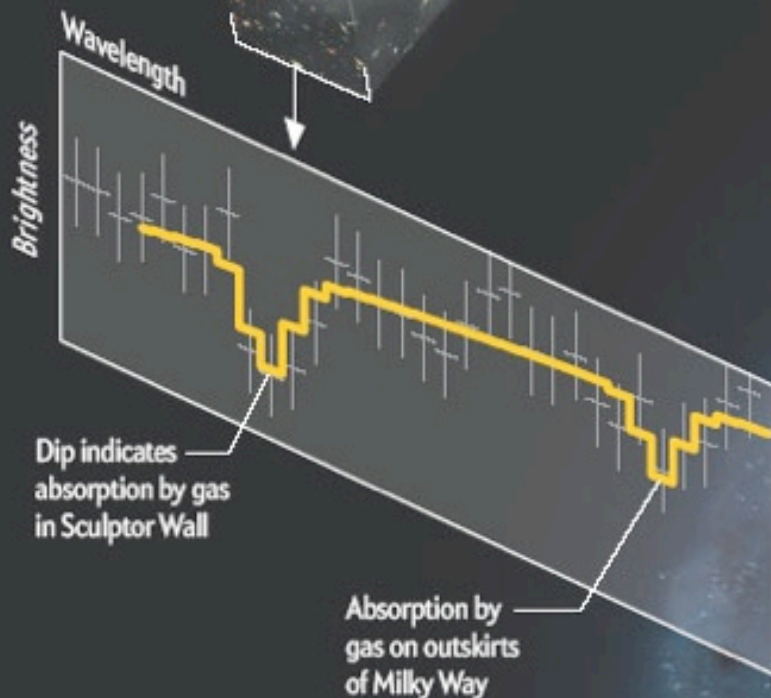
Component	Ω (ρ/ρ_c)
Dark Energy	0.691 ± 0.006
Matter (baryonic and non-baryonic)	0.312 ± 0.009
Baryons (Total)	0.0488 ± 0.0004
Baryons in stars and stellar remnants	~ 0.003
Neutrinos	~ 0.001
Photons (CMB)	5×10^{-5}

Betrayed by Its Shadow

Astronomers think they may have found where the bulk of the normal matter in the universe lurks: not in galaxies but in a form of intergalactic gas (mostly hydrogen) called the warm-hot intergalactic medium, or WHIM. The name connotes that the gas is less than blazingly hot and, consequently, glows too feebly to see directly. Looking in the interstices of a giant filament of galaxies called the Sculptor Wall, astronomers saw, in essence, the WHIM's shadow: the gas absorbed x-rays from a background object at a distinctive wavelength.

H 2356-309
(background
x-ray source)

Sculptor Wall



X-rays

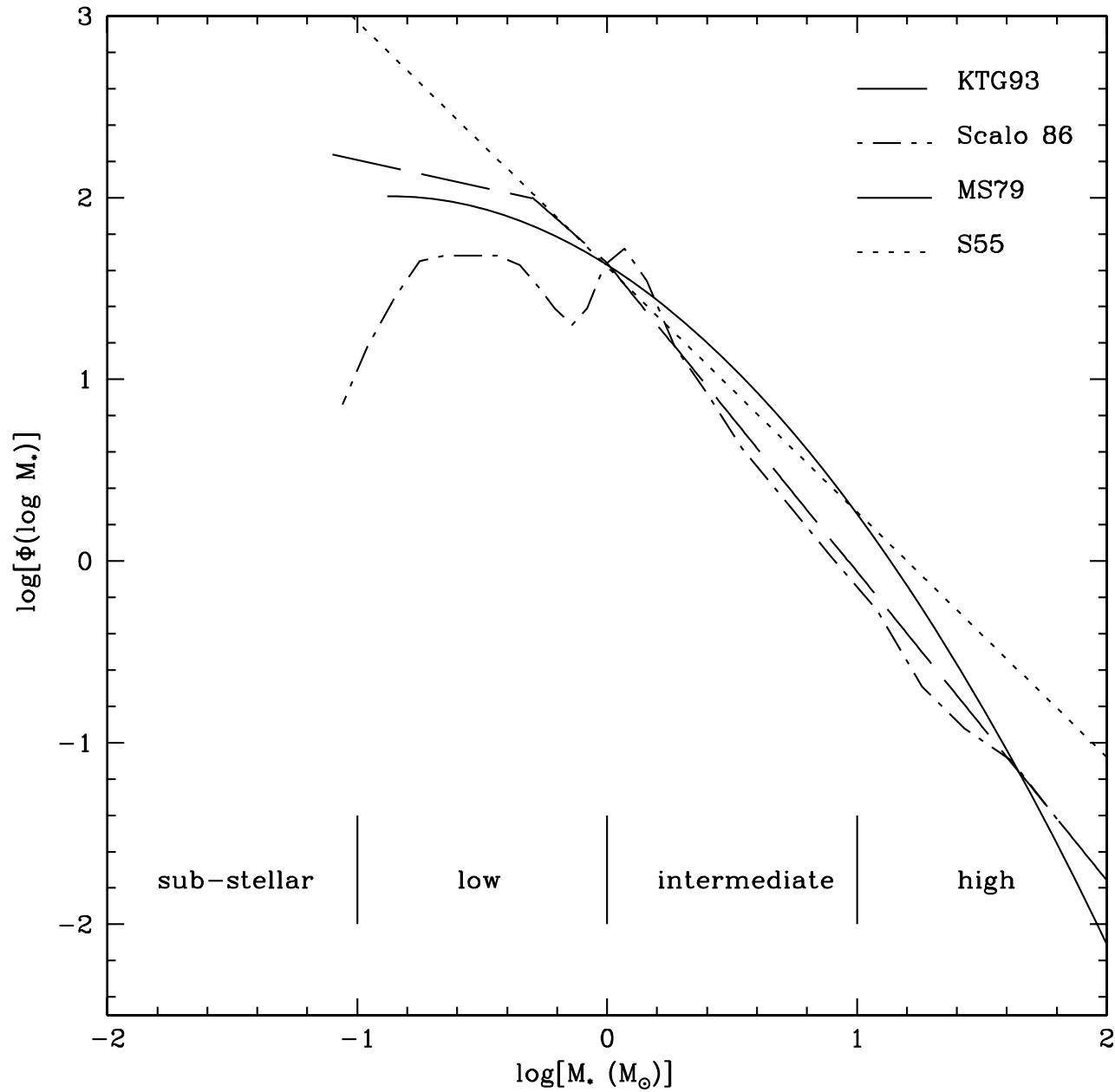
Chandra
X-ray Observatory

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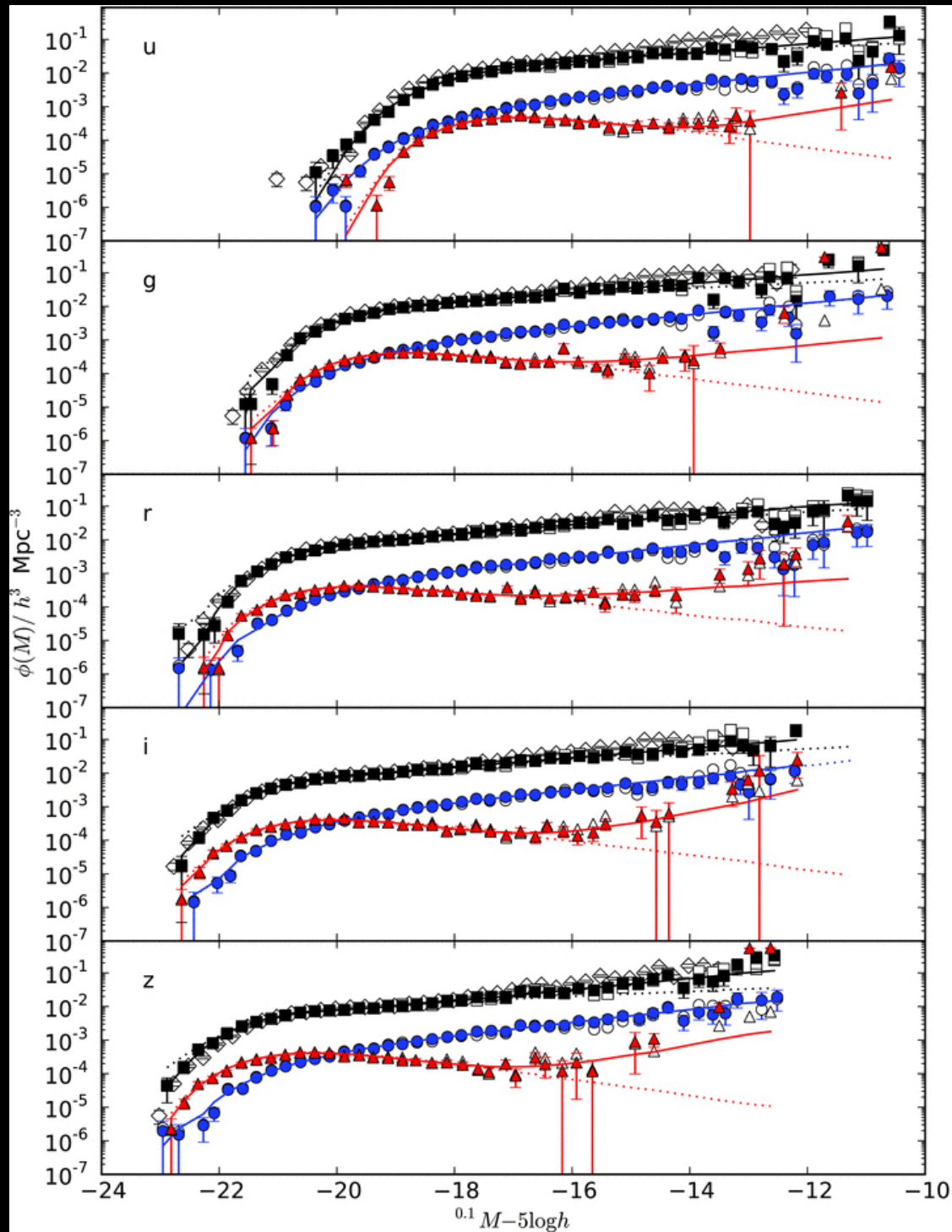
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Stellar Initial Mass Functions



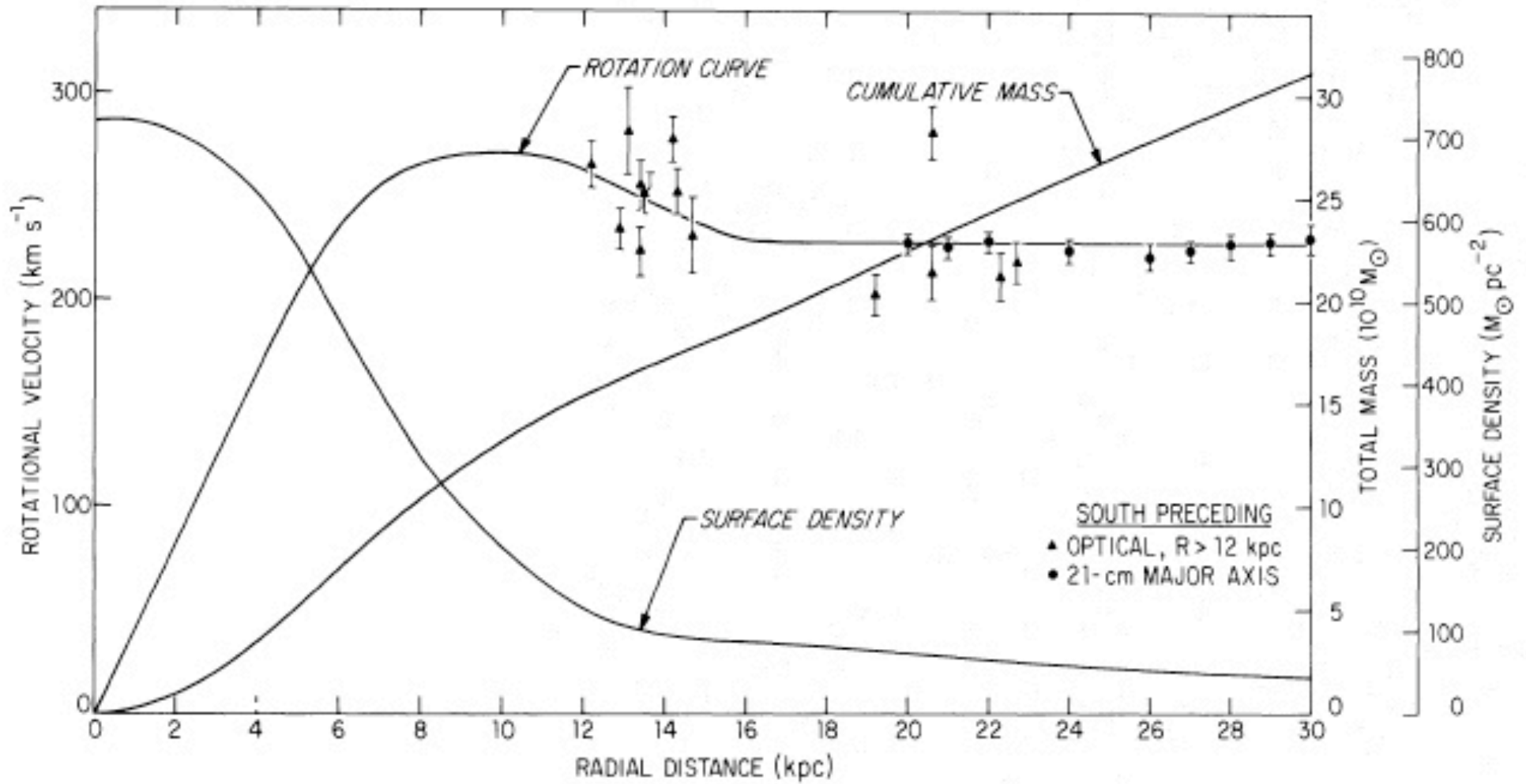
Meyer+ 1999

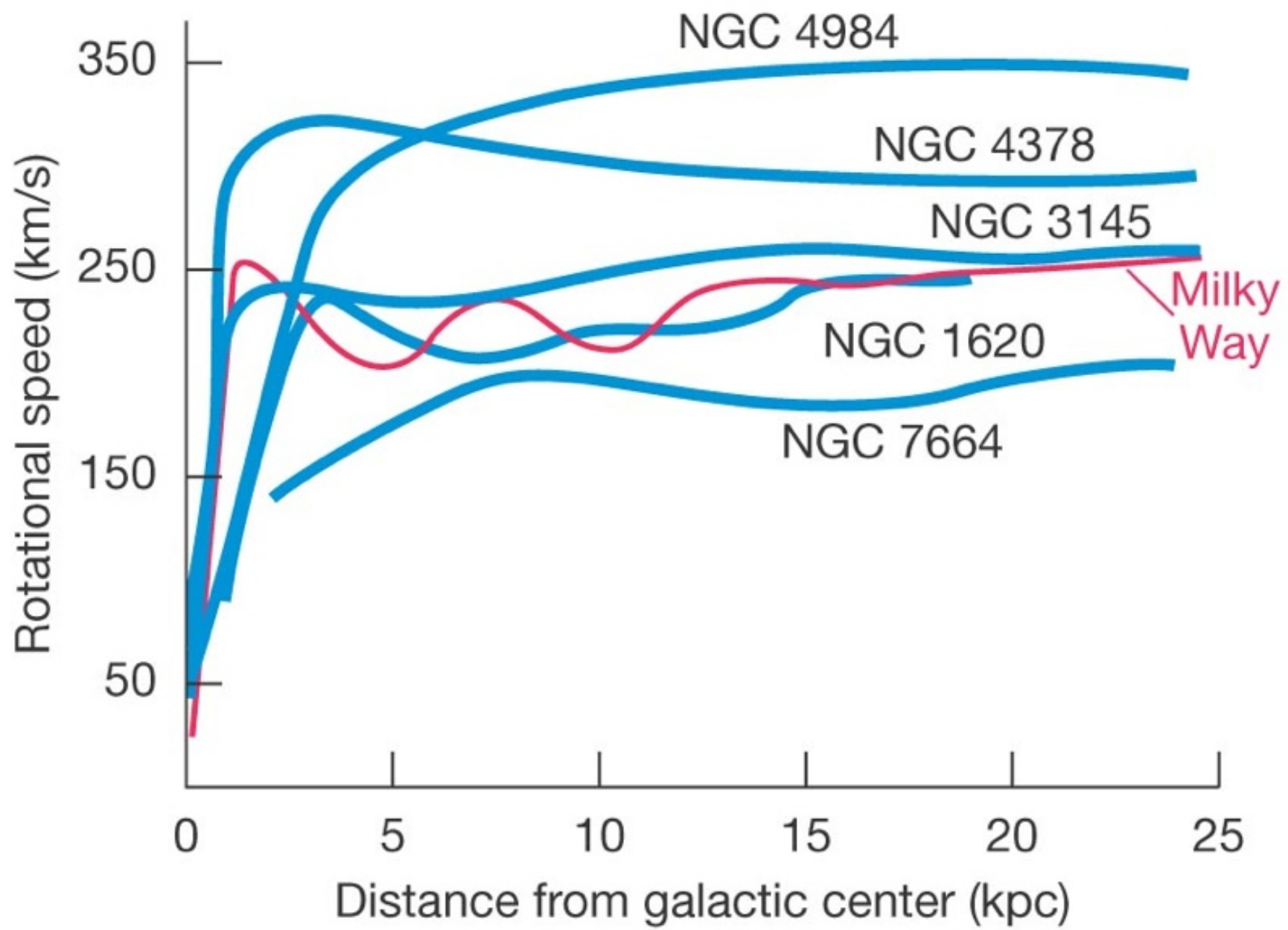


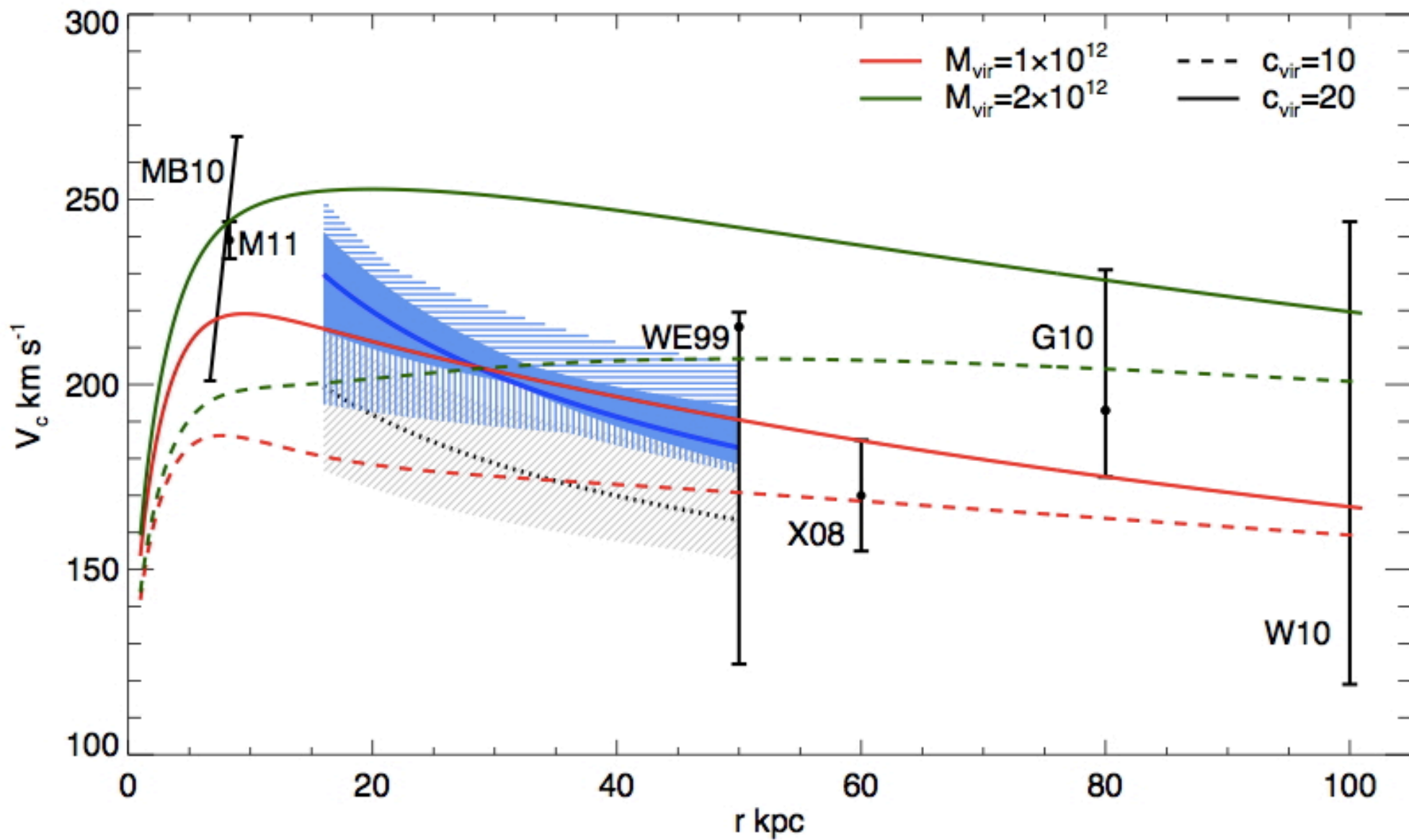
Loveday+ 2012





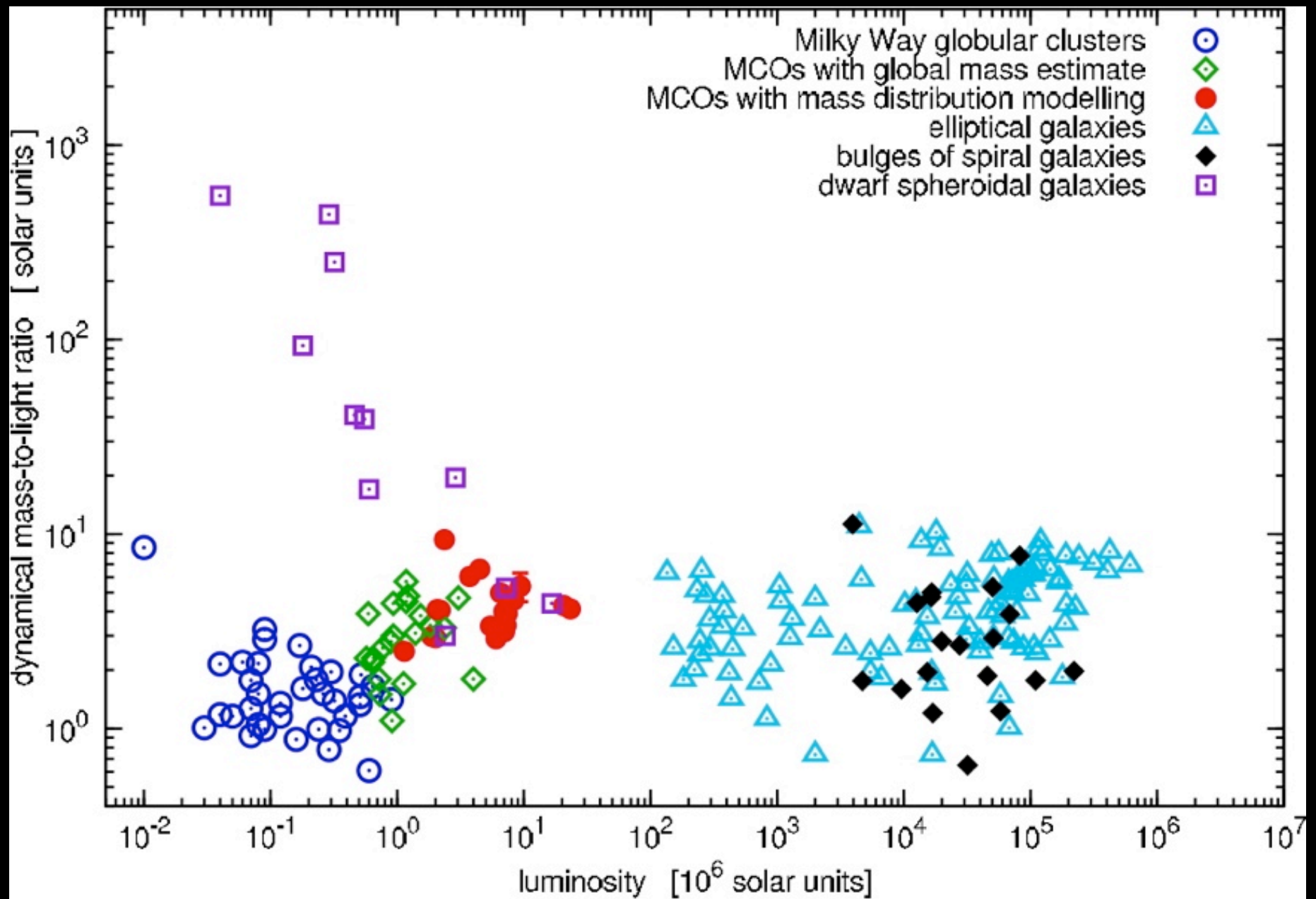






Deason+ 2012





Dabringhausen+ 2008



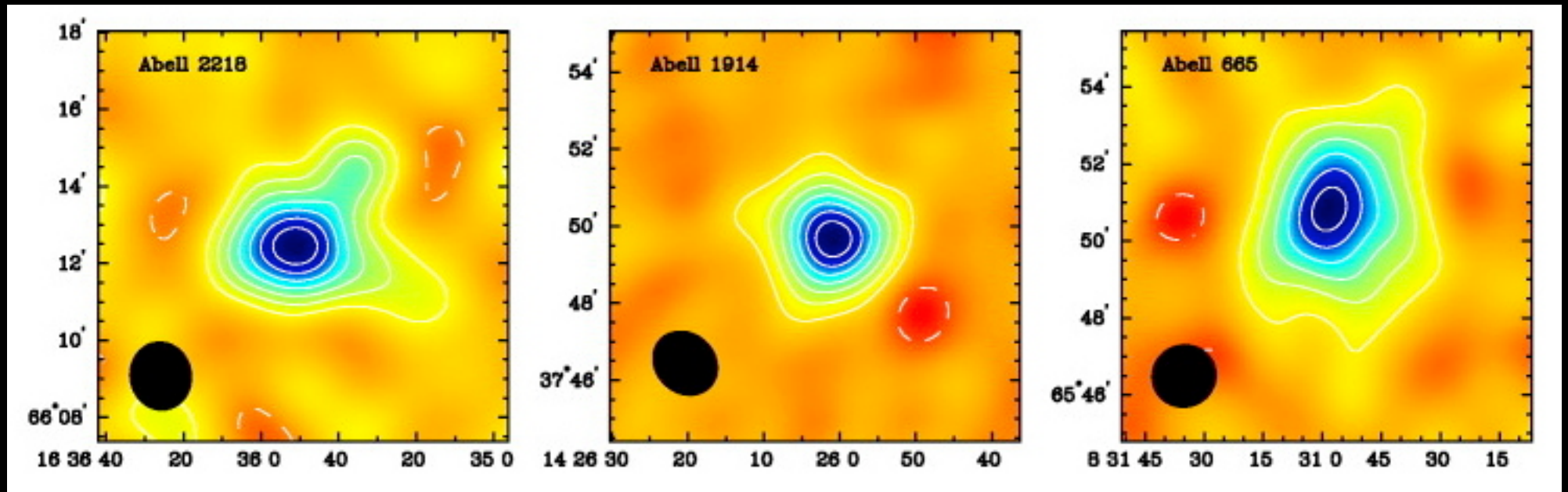


NASA, ESA, and the Hubble Heritage Team (STScI/AURA)

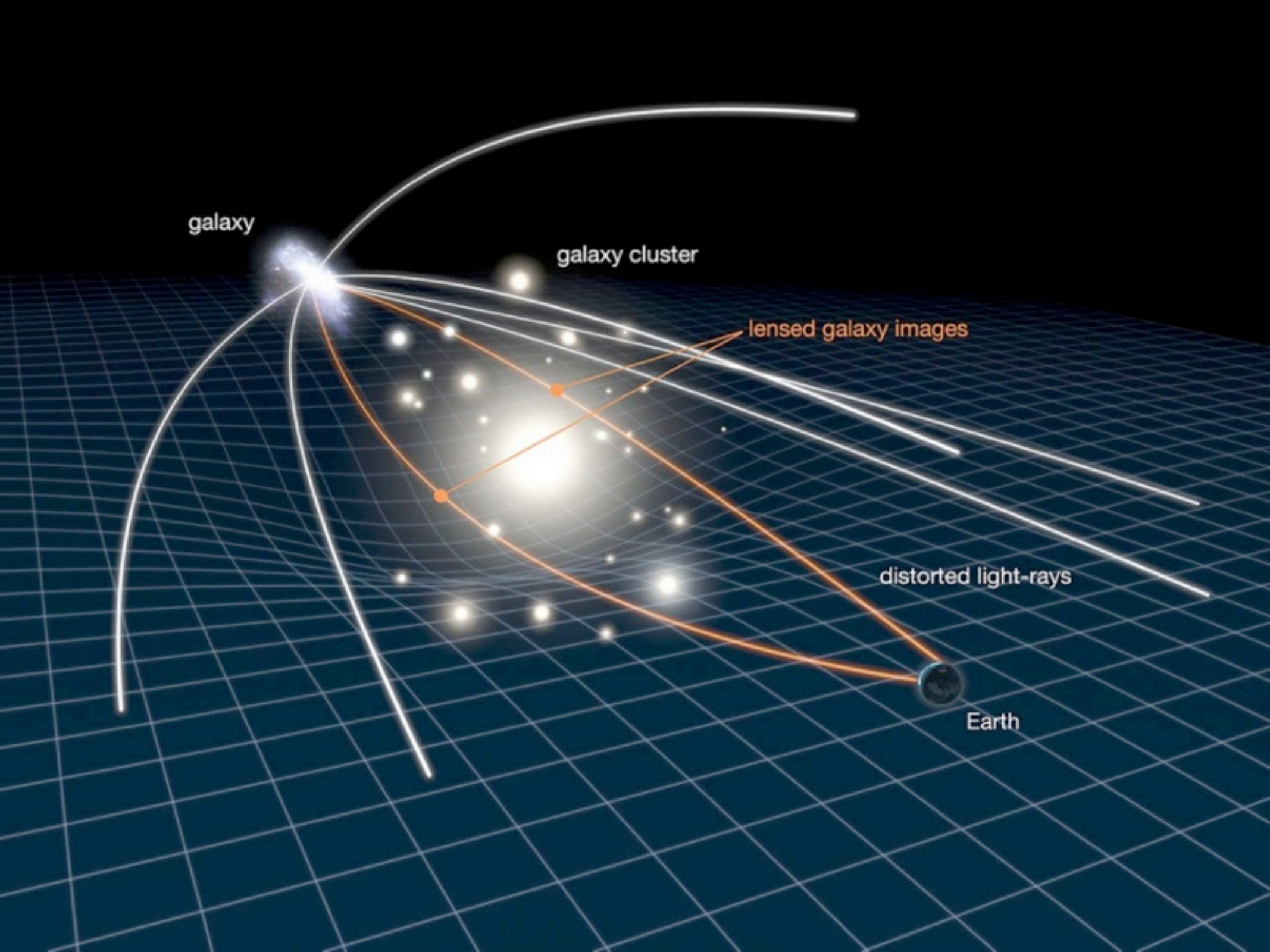








Grego+ 2001



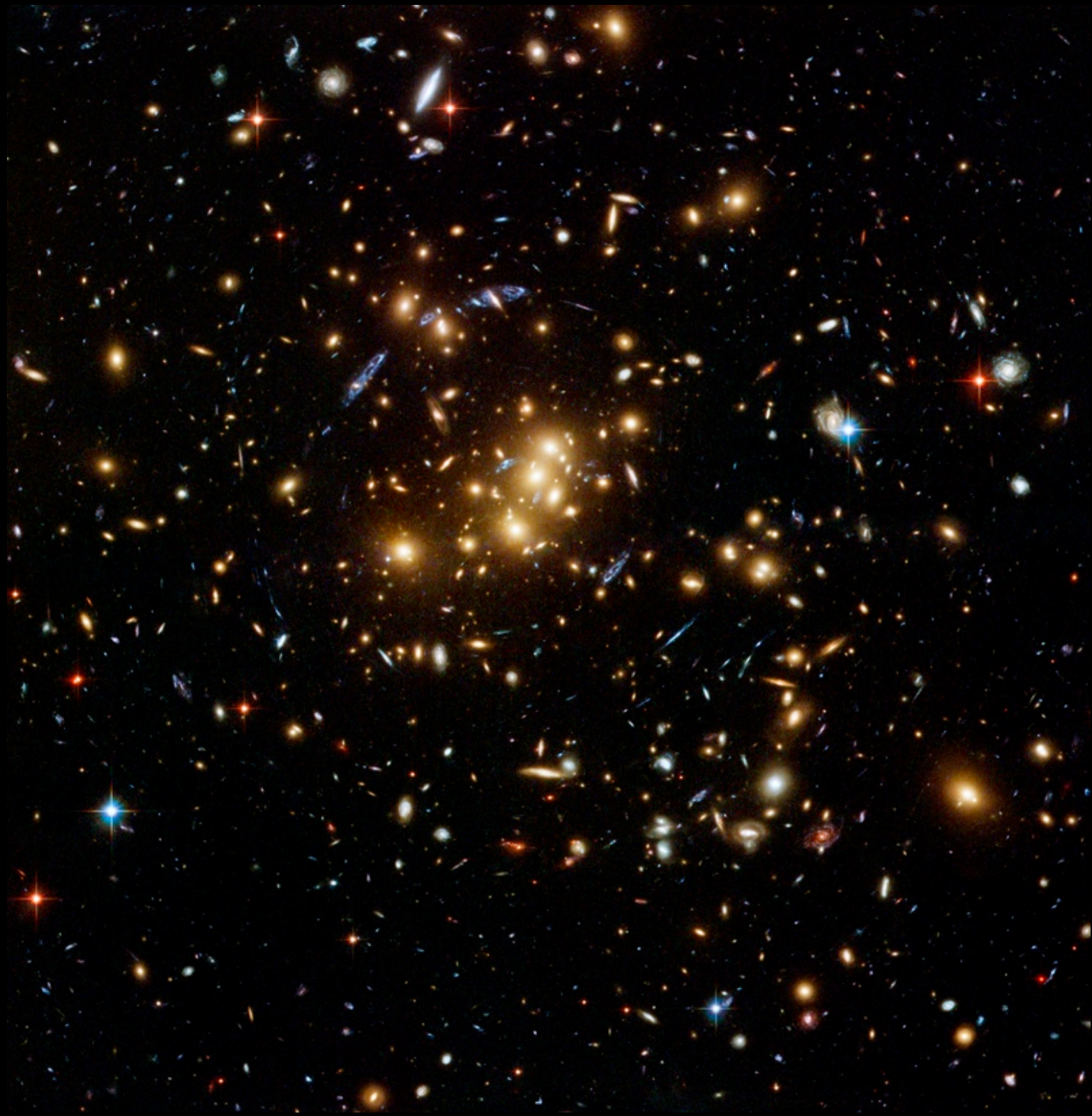
galaxy

galaxy cluster

lensed galaxy images

distorted light-rays

Earth





Gravitational lensing gives extra insight



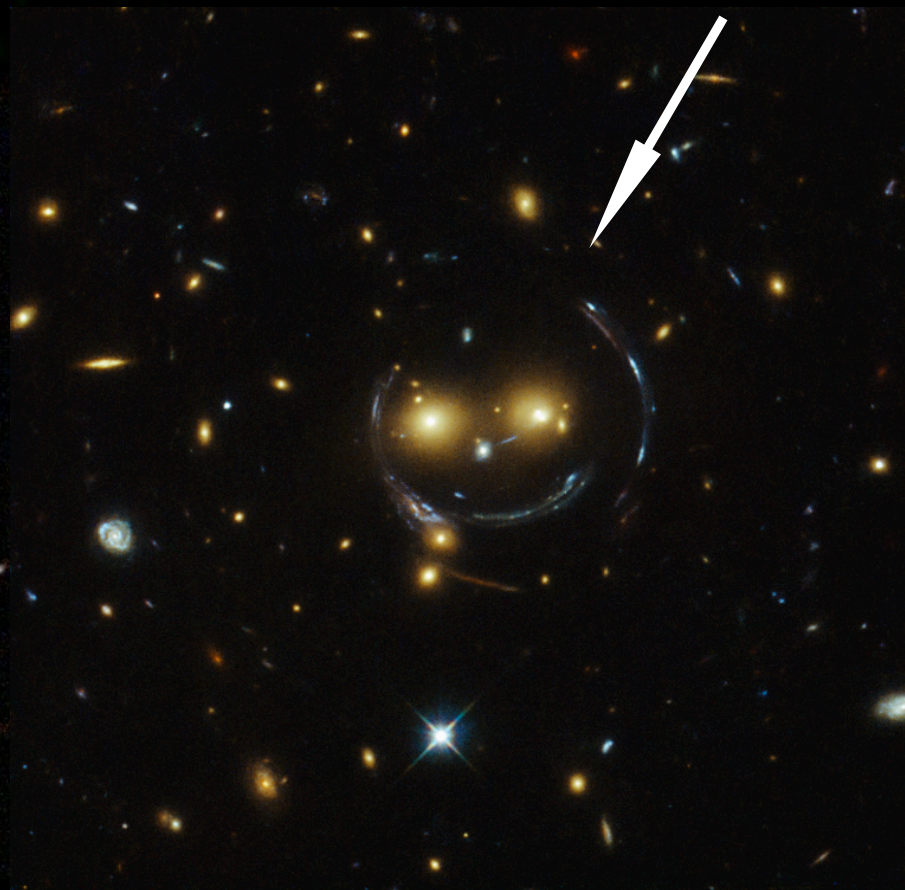
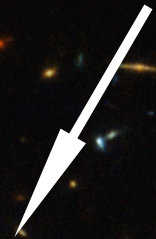
Cosmic Eye, $z = 3.073$



Cosmic Horseshoe, $z = 2.379$



Smiley, $z = 0.97$



DEFLECTION OF LIGHT RAYS CROSSING THE UNIVERSE, EMITTED BY DISTANT GALAXIES

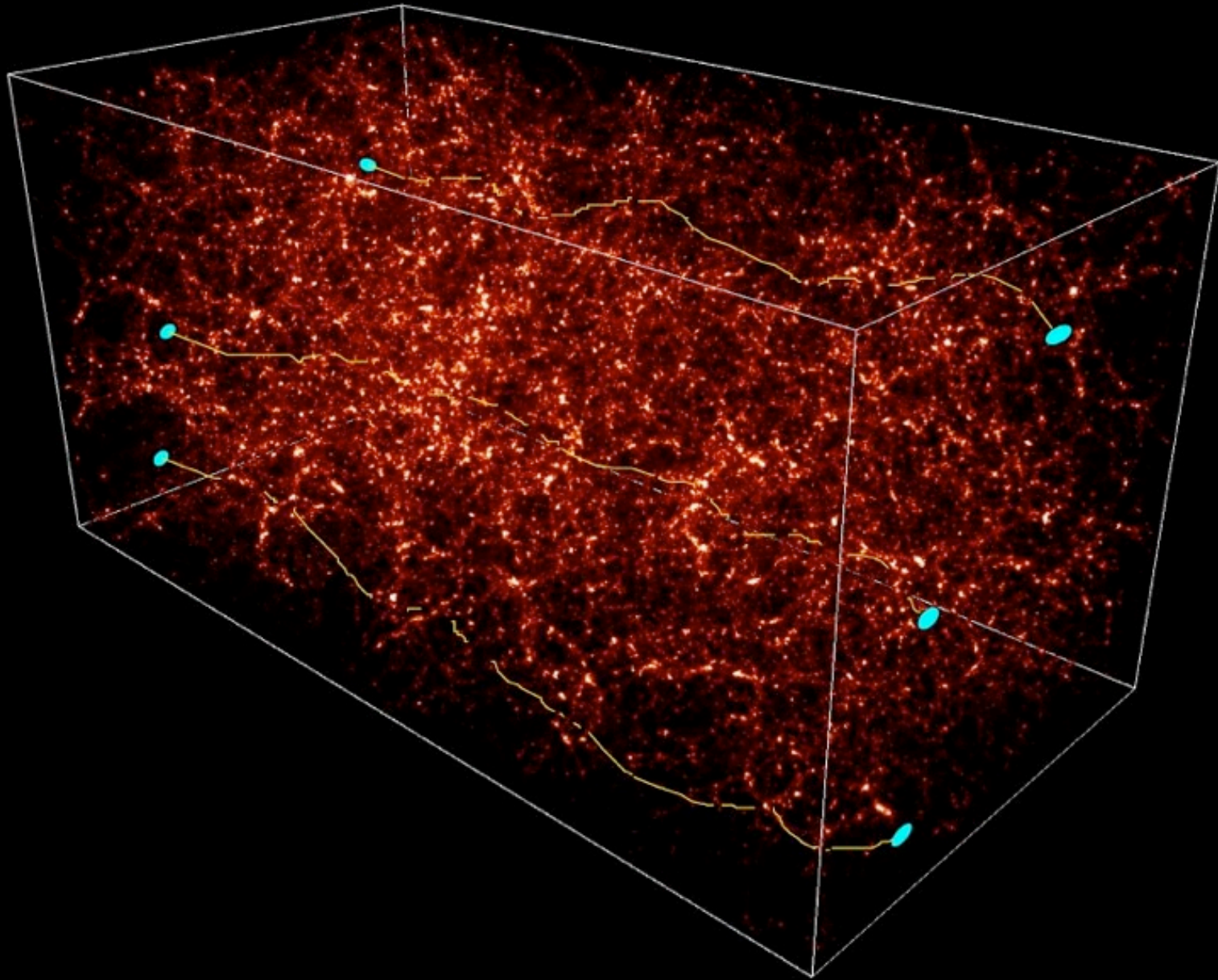
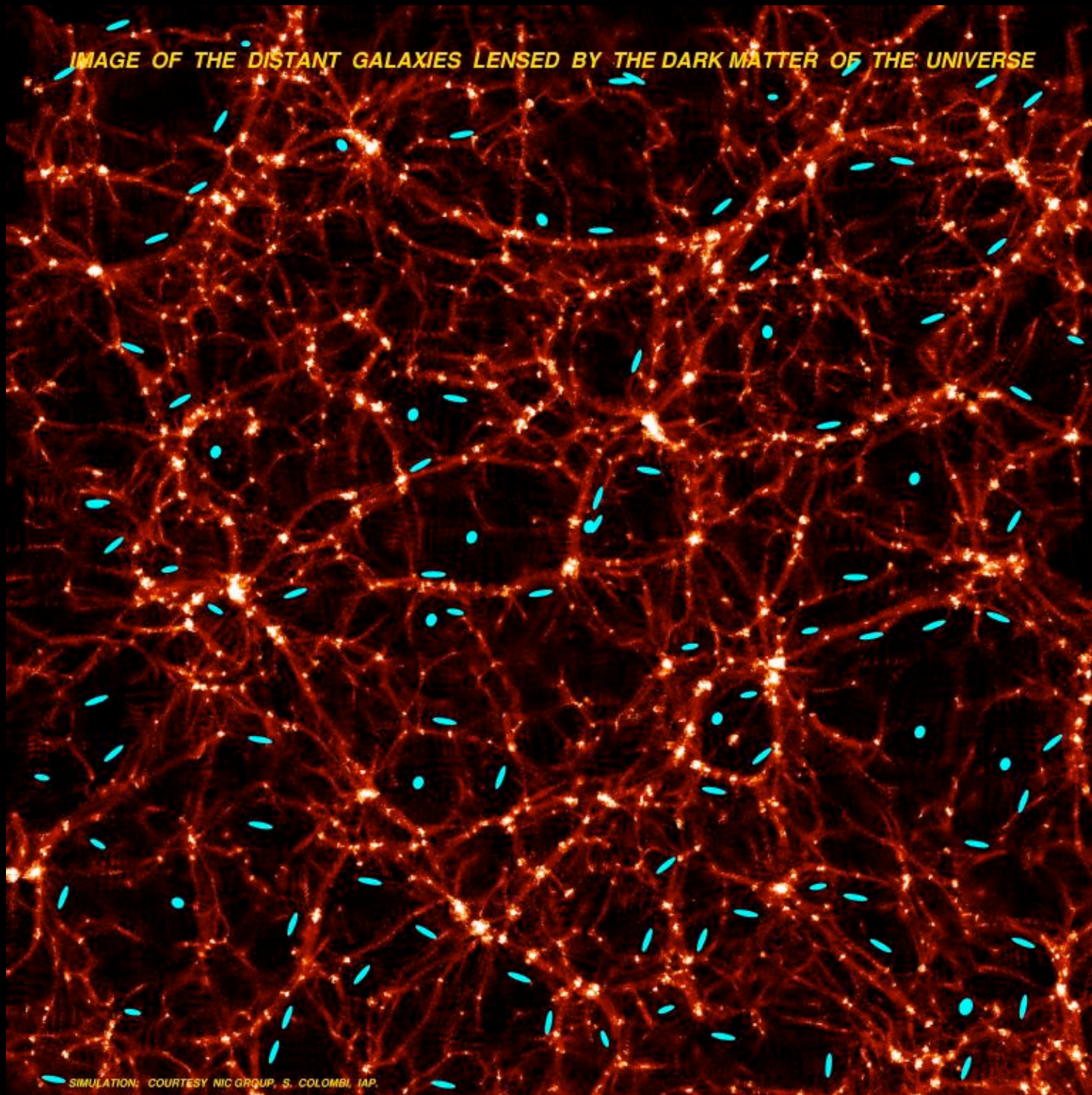
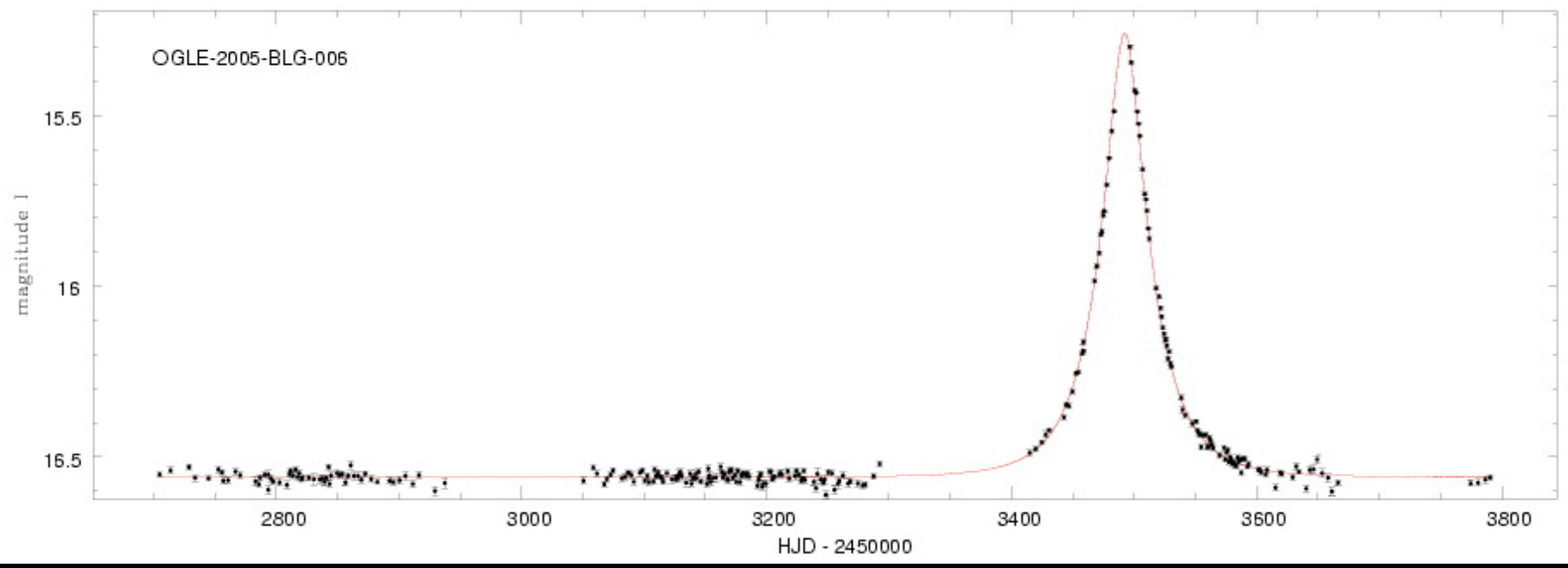


IMAGE OF THE DISTANT GALAXIES LENSED BY THE DARK MATTER OF THE UNIVERSE

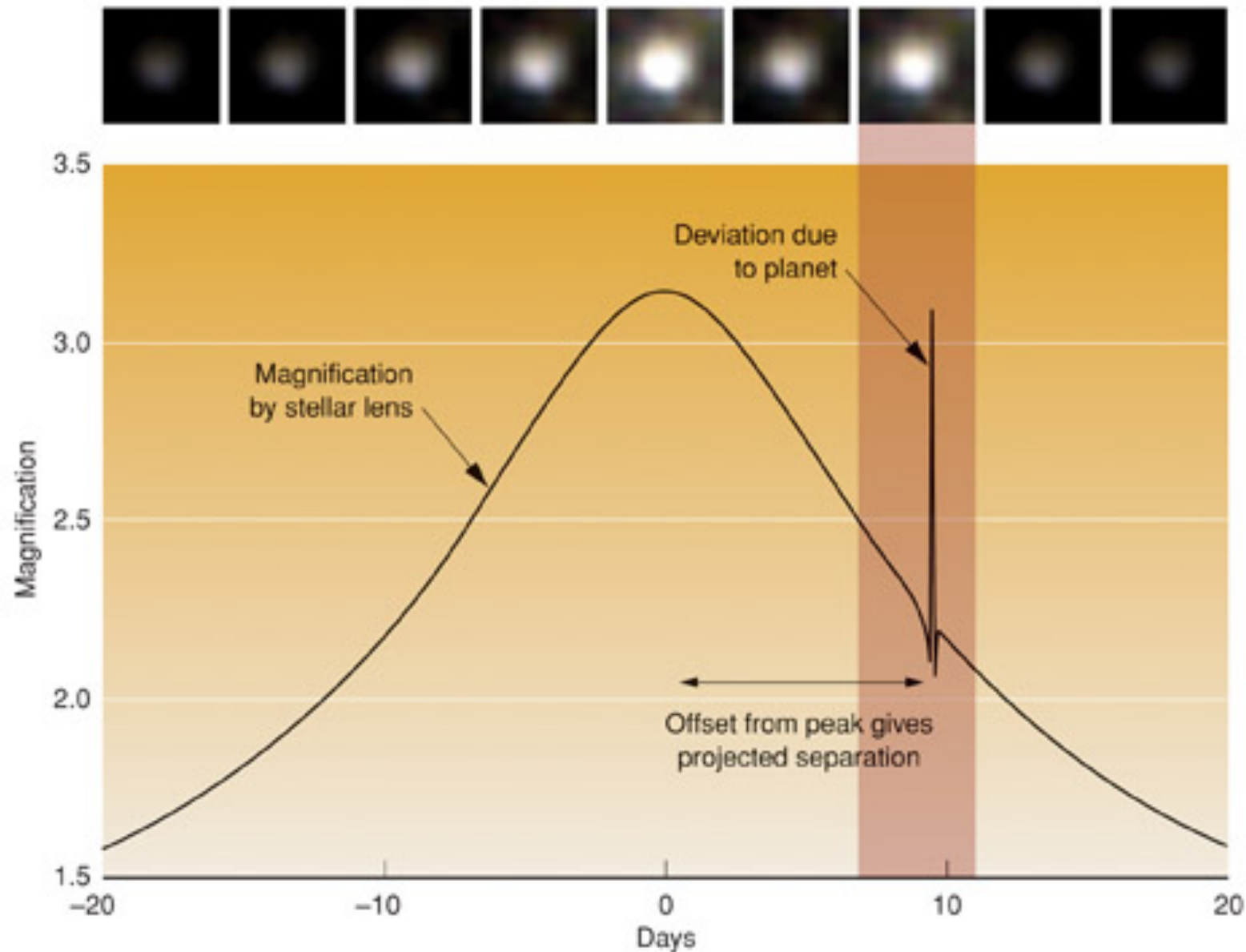


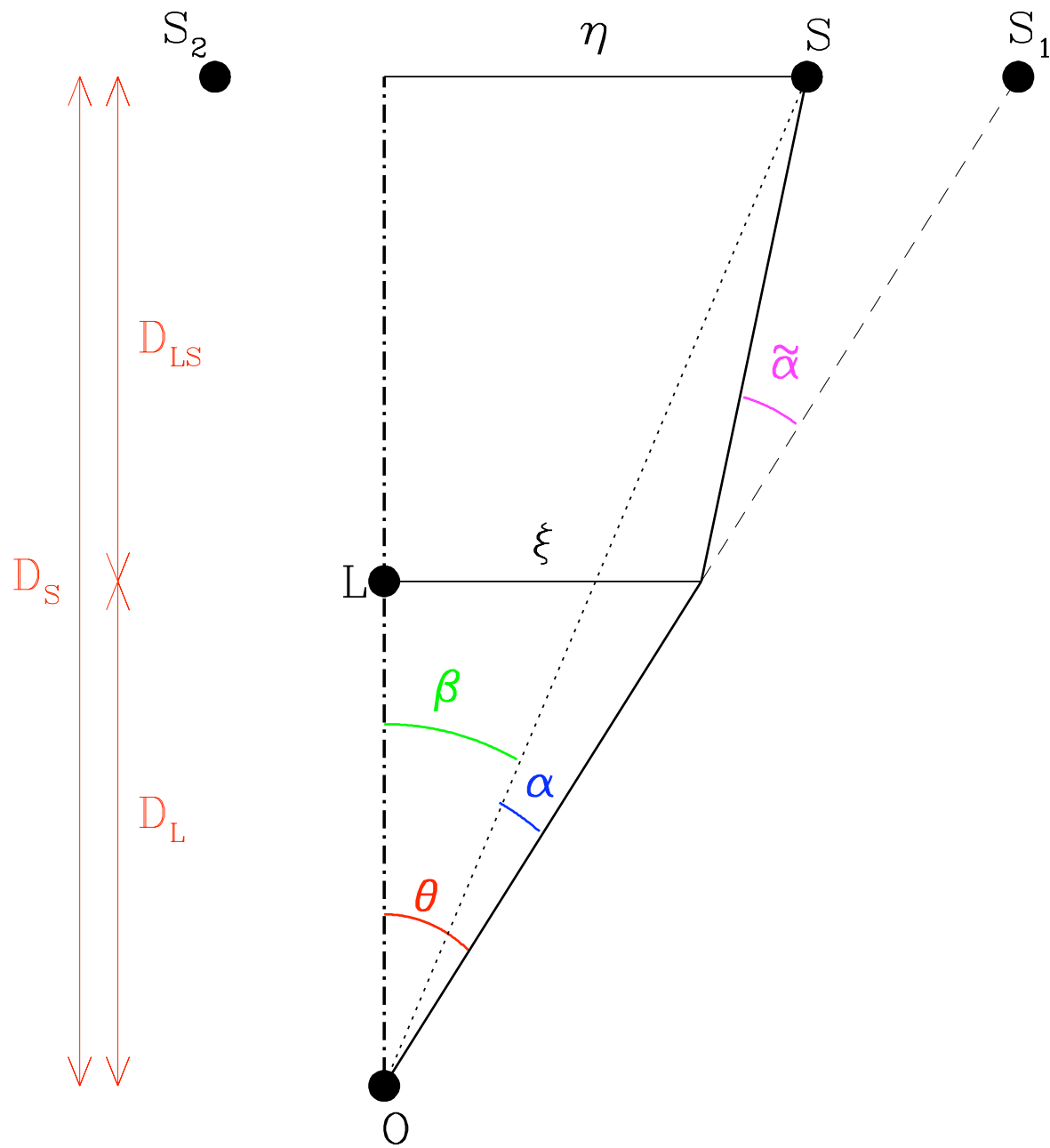
SIMULATION: COURTESY NIC GROUP, S. COLOMBI, IAP.



Planet searches

Star and planet act as microlenses





Gravitational lensing gives extra insight



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